

# Some observations about quantum chemistry software GAUSSIAN

## Abstract

When laboratory study of some molecule is not available, one may plan to use data obtained from Quantum Chemistry software, such as GAUSSIAN, MOLPRO, NWCHEM, etc. For our investigation of cosmic molecules, we need reliable data for rotational and centrifugal distortion constants. For some molecules, we have obtained these data with the help of Quantum Chemistry software GAUSSIAN and compared them with those obtained from the laboratory studies. We have found that in some cases, the two sets of data are very close to each other whereas in some cases, they differ very much. As the laboratory measurements provide the most reliable data, one would like to use the GAUSSIAN data only when the laboratory data are available. Thus, an obvious question arises how to decide the reliability of GAUSSIAN data, when for that particular molecule no laboratory data are available. Further, when the laboratory data are available, no one would like to use the GAUSSIAN data.

Volume 3 Issue 2 - 2019

**Mohit K Sharma**

Amity Centre for Astronomy &amp; Astrophysics, Amity Institute of Applied Sciences, Amity University, India

**Correspondence:** Mohit K Sharma, Amity Centre for Astronomy & Astrophysics, Amity Institute of Applied Sciences, Amity University, Noida 201313, India,  
 Email mksharma4@amity.edu, mohitkuamrsharma32@yahoo.in

**Received:** December 21, 2018 | **Published:** April 03, 2019

## Introduction

In a cosmic object having molecules, kinetic temperature in general is very low; few tens of Kelvin. Thus, one is concerned with the rotational levels in the ground vibrational state and ground electronic state. The rotational and centrifugal distortion constants, electric dipole moment can be used for calculation of energies of rotational levels and radiative transition probabilities (Einstein  $A$ -coefficients) for radiative transitions between the levels. We have investigated some molecules where laboratory data are available and for the same molecules we have obtained the data with the help of GAUSSIAN also. We have found that for some molecules, the two sets of data are in good agreement whereas for some molecules, they differ very much. As the laboratory data are the most reliable, one would like to use the GAUSSIAN data only in absence of the laboratory data. Thus, an obvious question arises how to decide the reliability of GAUSSIAN data. We are aware of the fact that the frequencies of spectral lines obtained from the GAUSSIAN data are not as accurate as required by the astronomers. However, the GAUSSIAN data can play important role in getting qualitative results about a molecule. We could not succeed in running the CCSD and CCSD (T) methods for the GAUSSIAN, as the computer program broke down each time during the execution. Therefore, we have employed the functional B3LYP method, i.e., Becke's three parameter exchange function B<sup>3</sup> (Becke<sup>1</sup>) with Lee, Yang and Parr's gradient corrected exchange-correlation functional.<sup>2</sup>

## Investigation

In the present discussion, we have considered the following molecules following sections. -to be deleted.

### Cyclopropanone

Guillemin et al.,<sup>3</sup> have recorded spectrum of cyclopropanone ( $c\text{-C}_3\text{H}_2\text{O}$ ) and have derived rotational and centrifugal distortion constants for Watson's rotational operator written in  $I'$  representation and - to be deleted with A-type reduction, given in Table 1 (column 2). Sharma et al.<sup>4</sup> have optimized the cyclopropanone with the help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> using B3LYP method and cc-pVDZ basis set. The values are given in Table 1 (column 3). The two sets of

data are in good agreement. The deviations of rotational constants  $A$ ,  $B$  and  $C$ , with respect to their experimental values are 0.46%, 1.18% and 0.90%, respectively.

**Table 1** Rotational and centrifugal distortion constants in MHz of  $c\text{-C}_3\text{H}_2\text{O}$

Constant	Laboratory	cc-pVDZ
A	32040.73	31894.85
B	7825.046	7733.81
C	6280.685	6224.503
$\Delta J$	$1.79362 \times 10^{-3}$	$1.651250766 \times 10^{-3}$
$\Delta JK$	$33.7882 \times 10^{-3}$	$3.271020188 \times 10^{-2}$
$\Delta K$	$50.65 \times 10^{-3}$	$4.379488485 \times 10^{-2}$
$\delta J$	$0.38536 \times 10^{-3}$	$3.533088298 \times 10^{-4}$
$\delta K$	$22.156 \times 10^{-3}$	$2.093510188 \times 10^{-2}$
$H_J$		$2.937852086 \times 10^{-10}$
$H_{JK}$		$6.851055671 \times 10^{-8}$
$H_{KJ}$		$-1.128055614 \times 10^{-7}$
$H_K$		$9.251718141 \times 10^{-8}$
$h_J$		$2.362038947 \times 10^{-10}$
$h_{JK}$		$3.968188988 \times 10^{-8}$
$h_K$		$1.205246347 \times 10^{-6}$

### Titanium dihydride

Inspired with good agreement between two sets of data for  $c\text{-C}_3\text{H}_2\text{O}$ , Sharma et al.,<sup>6</sup> decided to go for the investigation of titanium dihydride ( $\text{TiH}_2$ ) for which laboratory data are not available. They<sup>6</sup> have optimized the molecule  $\text{TiH}_2$  with the help of GAUSSIAN 2009 Frisch et al.,<sup>5</sup> where B3LYP method and cc-pVTZ basis set are used. The rotational and centrifugal distortion constants obtained for Watson's rotational operator written in  $I'$  representation and - to be deleted with A-type reduction are given in Table 2. The outcome of the investigation is very exciting.

**Table 2** Rotational and centrifugal distortion constants in MHz of TiH<sub>2</sub>

Constant	cc-pVTZ	Constant	cc-pVTZ
A	$2.8589602 \times 10^5$	$\Phi_J$	$2.037108491 \times 10^{-3}$
B	$1.2520818 \times 10^5$	$\Phi_{JK}$	$-1.928681064 \times 10^{-2}$
C	$8.707408 \times 10^4$	$\Phi_{KJ}$	$-1.946479675 \times 10^{-2}$
$\Delta J$	6.004343	$\Phi_K$	$5.110638046 \times 10^{-1}$
$\Delta JK$	$-4.186548023 \times 10^1$	$\varphi_J$	$1.014000418 \times 10^{-3}$
$\Delta K$	$2.549058853 \times 10^2$	$\varphi_{JK}$	$-3.557138924 \times 10^{-3}$
$\delta J$	2.465123	$\varphi_K$	$1.085768474 \times 10^{-1}$
$\delta K$	2.332384		

### Ethylene oxide

Pan et al.<sup>7</sup> recorded spectrum of ethylene oxide (*c*-C<sub>2</sub>H<sub>4</sub>O) and have derived rotational and centrifugal distortion constants for Watson's rotational operator written in *I'* representation and - to be deleted with A-type reduction, given in Table 3 (column 2). Sharma et al.<sup>8</sup> have optimized the ethylene oxide with the help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> bf using B3LYP method and cc-pVDZ basis set. The values of data are given in Table 3 (column 3). The two sets of data are in good agreement. The deviations of rotational constants *A*, *B* and *C*, with respect to their experimental values are -0.79%, -0.58% and -0.70%, respectively. It again provided us a confidence about our investigation of TiH<sub>2</sub> Sharma et al.<sup>6</sup>

### Vinylidene

Inspired with good agreement between two sets of data for *c*-C<sub>2</sub>H<sub>2</sub>O and for *c*-C<sub>2</sub>H<sub>4</sub>O, we decided to go for the investigation of vinylidene (H<sub>2</sub>CC) for which also laboratory data are not available. bf Sharma et al.<sup>9</sup> have optimized the molecule H<sub>2</sub>CC with the help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> employing the B3LYP method in conjunction with four basis sets, cc-pVTZ, aug-cc-pVDZ, aug-cc-pVTZ and aug-cc-pVQZ. The resulting rotational and centrifugal distortion constants for Watson's rotational operator written in *I'* representation and - to be deleted with S-type reduction are given in Table 4. There is good agreement between the data obtained from different basis sets.

### Silanone

Bailleux et al.<sup>10</sup> have recorded spectrum of silanone (H<sub>2</sub>SiO) and derived rotational and centrifugal distortion constants for Watson's rotational operator written in *I'* representation and - to be deleted with A-type reduction, given in Table 5 (column 2). Sharma et al.<sup>11</sup> have optimized the molecule H<sub>2</sub>SiO with the help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> bf employing B3LYP method in conjunction with three basis sets, aug-cc-pVDZ, aug-cc-pVTZ and aug-cc-pVQZ. The values are given in Table 5 (columns 3-5). There is very good agreement between the four sets of data. The deviations of rotational constants *A*, *B* and *C* obtained for the basis set aug-cc-pVQZ, with respect to their experimental values are 0.24%, 0.53% and 0.18%, respectively.

### cis-Formic acid

Winnerwiser et al.<sup>12</sup> have recorded spectrum of cis-Formic acid (cis-HCOOH) and derived rotational and centrifugal distortion constants for Watson's rotational operator written in *I'* representation and - to be deleted with A-type reduction, given in Table 6 (column 2). Sharma et al.<sup>13</sup> have optimized the molecule cis-HCOOH with the

help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> bf employing B3LYP method in conjunction with three basis sets, aug-cc-pVDZ, aug-cc-pVTZ and aug-cc-pVQZ. The values are given in Table 6 (columns 3-5). There is very good agreement between the four sets of data. The deviations of rotational constants *A*, *B* and *C* obtained for the basis set aug-cc-pVQZ, with respect to their experimental values are -1.16%, -0.22% and -0.46%, respectively. All these data provided us a encouragement about our investigations of TiH<sub>2</sub> and H<sub>2</sub>CC molecules.

**Table 3** Rotational and centrifugal distortion constants (MHz) of *c*-C<sub>2</sub>H<sub>4</sub>O

Constant	Laboratory	cc-pVDZ
A	25483.89	25685.72
B	22120.85	22249.95
C	14097.84	14197.62
$\Delta J$	$51.1883 \times 10^{-3}$	$50.79096319 \times 10^{-3}$
$\Delta JK$	$-70.4938 \times 10^{-3}$	$-71.09013012 \times 10^{-3}$
$\Delta K$	$27.6541 \times 10^{-3}$	$28.41532834 \times 10^{-3}$
$\delta J$	$-9.01689 \times 10^{-3}$	$8.836730565 \times 10^{-3}$
$\delta K$	$3.3491 \times 10^{-3}$	$-6.556467036 \times 10^{-3}$
$\Phi_J$	$0.2456 \times 10^{-6}$	$-5.960452772 \times 10^{-8}$
$\Phi_{JK}$	$-5.2164 \times 10^{-6}$	$-4.838170644 \times 10^{-6}$
$\Phi_{KJ}$	$15.7370 \times 10^{-6}$	$15.30527314 \times 10^{-6}$
$\Phi_K$	$-10.638 \times 10^{-6}$	$-10.40655085 \times 10^{-6}$
$\varphi_J$	$-0.05097 \times 10^{-6}$	$-2.887999339 \times 10^{-8}$
$\varphi_{JK}$	$1.4297 \times 10^{-6}$	$-1.411514955 \times 10^{-6}$
$\varphi_K$	$-17.8633 \times 10^{-6}$	$1.713117251 \times 10^{-5}$
L <sub>J</sub>	$-0.1210 \times 10^{-9}$	
L <sub>JJK</sub>	$-0.1288 \times 10^{-9}$	
L <sub>JK</sub>	$0.624 \times 10^{-9}$	
L <sub>KKJ</sub>	$-0.800 \times 10^{-9}$	
L <sub>K</sub>	$0.892 \times 10^{-9}$	
I <sub>J</sub>	$-0.00367 \times 10^{-9}$	
I <sub>JK</sub>	$0.0921 \times 10^{-9}$	
I <sub>KJ</sub>	$-0.448 \times 10^{-9}$	
I <sub>K</sub>	$0.679 \times 10^{-9}$	
P <sub>K</sub>	$-1.114 \times 10^{-12}$	

### Disilicon

McCarthy et al.<sup>14</sup> has recorded spectrum of disilicon (Si<sub>2</sub>C) and have derived rotational and centrifugal distortion constants for Watson's rotational operator written in *I'* representation and - to be deleted with S-type reduction, given in Table 7 (column 2). Sharma et al.<sup>15</sup> have optimized the molecule Si<sub>2</sub>C with the help of GAUSSIAN 2009 Frisch et al.<sup>5</sup> employing the B3LYP method in conjunction with three basis sets, aug-cc-pVDZ, aug-cc-pVTZ and aug-cc-pVQZ. The values are given in Table 7 (columns 3-5). There is large disagreement between the laboratory data and those obtained from GAUSSIAN. The deviations of rotational constants *A*, *B* and *C* obtained for the basis set aug-cc-pVQZ, with respect to their experimental values are -51.89%, 20.23% and 15.28%, respectively. These large deviations perturbed us and lead to a question about the reliability of GAUSSIAN data.

**Table 4** Rotational and centrifugal distortion constants in MHz of H<sub>2</sub>CC

Constant	cc-pVTZ	aug-cc-pVDZ	aug-cc-pVTZ	aug-cc-pVQZ
A × 10 <sup>-5</sup>	2.858903	2.858903	2.858903	2.858903
B × 10 <sup>-4</sup>	3.993411	3.993411	3.993411	3.993411
C × 10 <sup>-4</sup>	3.503965	3.503965	3.503965	3.503965
D <sub>J</sub> × 10 <sup>3</sup>	42.70791	43.44087	44.28356	44.55227
D <sub>JK</sub>	20.98798	18.94745	20.44459	20.4197
D <sub>K</sub>	5.317725	6.440168	5.879314	5.933398
d <sub>1</sub> × 10 <sup>2</sup>	-1.442126677	-1.373607832	-1.445080323	-1.452037644
d <sup>2</sup> × 10 <sup>2</sup>	-2.675152631	-2.408896481	-2.609261594	-2.613141728
H <sub>J</sub> × 10 <sup>6</sup>	-9.085110924	-7.573403608	-8.601743486	-8.655849617
H <sub>JK</sub> × 10 <sup>3</sup>	2.717971	2.280665	2.564823	2.580959
H <sub>KJ</sub> × 10 <sup>2</sup>	-1.854934803	-1.597951879	-1.709596957	-1.731162015
H <sub>K</sub> × 10 <sup>2</sup>	2.401038	2.151629	2.272677	2.29427
h <sub>1</sub> × 10 <sup>6</sup>	-1.524539329	-1.248551723	-1.444404960	-1.443374339
h <sub>2</sub> × 10 <sup>6</sup>	4.698862	3.929762	4.456302	4.485417
h <sub>3</sub> × 10 <sup>6</sup>	1.651997	1.367133	1.5717	1.57211

**Table 5** Rotational and centrifugal distortion constants in MHz of H<sub>2</sub>SiO

Constant	Experiment	cc-pVDZ	aug-cc-pVDZ	aug-cc-pVTZ	aug-cc-pVQZ
A × 10 <sup>5</sup>	1.666573	1.62688	1.62316	1.658067	1.662584
B × 10 <sup>-4</sup>	1.867939	1.787253	1.78186	1.842979	1.858002
C × 10 <sup>-4</sup>	1.674277	1.610344	1.605601	1.65862	1.671235
ΔJ × 10 <sup>2</sup>	1.75216	1.631609	1.647716	1.66202	1.676962
ΔJK × 10 <sup>1</sup>	6.02486	5.543688	5.632878	5.839659	5.871728
ΔK	7.5	8.199677	8.090008	8.443277	8374.246
δJ × 10 <sup>3</sup>	2.0811	1.822749	1.836778	1.876133	1.906645
δK × 10 <sup>1</sup>	4.13	3.503961	3.550486	3.660305	3.689542
ΦJ × 10 <sup>9</sup>		6.733706	6.296932	5.253595	5.657116
ΦJK × 10 <sup>6</sup>	4.757	7.800663	7.879044	8.266248	8.41835
ΦKJ × 10 <sup>5</sup>	-4.774	-1.57875	-1.796	-2.14959	-2.2458
ΦK × 10 <sup>3</sup>		1.443688	1.377452	1.435596	1.415752
φJ × 10 <sup>9</sup>		3.572809	3.531365	3.408102	3.524366
φJK × 10 <sup>6</sup>		4.031514	4.065714	4.256998	4.334719
φK × 10 <sup>4</sup>		4.511626	4.605502	4.727325	4.732419

**Table 6** Rotational and centrifugal distortion constants (MHz) of cis-HCOOH

Constant	Lab	Optimization		
		aug-cc-pVDZ	aug-cc-pVTZ	aug-cc-pVQZ
A	86461.62	85967.44	87387.25	87478.32
B	11689.18	11617.78	11690.94	11715.35
C	10284	10234.65	10311.44	10331.7
$\Delta J \times 10^3$	8.35515	6.594446	6.564108	6.564716
$\Delta JK \times 10^3$	-71.4412	259.8124	274.8887	275.2381
$K \times 10^3$	2361.672	2176.623	2296.741	2303.991
$\delta J \times 10^3$	1.41773	0.592251	0.556347	0.55381
$\delta K \times 10^3$	40.747	115.2922	118.6733	118.4331
$\Phi J \times 10^8$	1.064	-0.4527703926	-0.4722888293	-0.4808865771
$\Phi JK \times 10^6$	-0.2974	9.566051	10.98867	11.10552
$\Phi KJ \times 10^6$	-9.673	108.8903	119.0774	119.4857
$\Phi K \times 10^6$	185.11	1.389687	1.452612	1.434686
$\varphi J \times 10^9$	2.317	0.108608	0.08907	0.084311
$\varphi JK \times 10^6$	-0.73	0.746964	0.783595	0.775275
$\varphi K \times 10^6$		37.1677	39.04597	38.76245
$L_K \times 10^9$	-20.2			
$I_{JK} \times 10^9$	0.558			

**Table 7** Rotational and centrifugal distortion constants of Si<sub>2</sub>C in MHz

Constant	Experimental	aug-cc-pVDZ	aug-cc-pVTZ	aug-cc-pVQZ
A	64074.34	115272.4	141935.4	133191.5
B	4395.621	3648.742	3597.055	3655.86
C	4102.028	3536.791	3508.148	3558.194
$D_J \times 10^3$	9.7315	18.02061	28.55172	17.97574
$D_{JK}$	-0.8572075	-7.016988202	-17.432464312	-9.250582631
$D_K \times 10^{-2}$	0.235881	7.258966	27.69195	12.60339
$d_1 \times 10^3$	1.519832	2.2482	3.425177	2.169344
$d_2 \times 10^1$	0.51591	1.269923	2.405429	1.420247
$H_J \times 10^7$	-0.41349	-6.979665512	-26.90729235	-4.538089470
$H_{JK} \times 10^4$	0.93298	9.032487	48.46837	10.80401
$H_{KJ} \times 10^1$	-0.0188755	-2.856368246	-22.85802143	-4.878977881
$H_K$	0.044863	26.42362	320.7195	61.69333
$h_j \times 10^8$	-0.5231	-11.16426814	-43.12215486	-6.341531159
$h_k \times 10^3$		1.388191	7.346143	3.083056
$h_{jk} \times 10^5$	-0.6586	-2.147381766	-9.373151002	-1.964934254

## Discussion

When got good agreement between the laboratory data and those obtained with the help of GAUSSIAN for  $c\text{-C}_3\text{H}_2\text{O}$ ,  $c\text{-C}_2\text{H}_4\text{O}$ ,  $\text{H}_2\text{SiO}$  and  $\text{cis-HCOOH}$ , we felt encouraged that in absence of laboratory data for a particular molecule, at least qualitative analysis of the molecule could be done with the help of the GAUSSIAN data. But, a large disagreement between the two sets of data for  $\text{Si}_2\text{C}$  has shattered down all the confidence. Thus, an obvious question arises how to decide the reliability of GAUSSIAN data, when laboratory data are not available. Further, when the laboratory data are available, no one would like to use the GAUSSIAN data. We Sharma et al.<sup>16</sup> have earlier presented some observations about the Quantum Chemistry software MOLPRO. Werner et al.<sup>17</sup> About the computer code MOLSCAT Hutson et al.,<sup>18</sup> we Sharma et al.<sup>19</sup> have presented some observations. About the observations, someone may respond - to be deleted These observations however - to be deleted provide some awareness to the users of the GAUSSIAN, MOLPRO and MOLSCAT.

## Acknowledgments

Author is grateful to Hon'ble Dr. Ashok K. Chauhan, Founder President, Hon'ble Dr. Atul Chauhan, Chancellor, and Hon'ble Vice Chancellor Dr. Balvinder Shukla, Amity University for valuable support and encouragements. He is thankful to the SERB, DST, New Delhi for awarding the NPDP.

## Conflicts of interest

The author declares there is no conflict of interest.

## References

1. Becke AD. Density-functional thermochemistry. III. The role of exact exchange. *The Journal of Chemical Physics*. 1993;98:5648–5652.
2. Lee C, Yang W, Parr RG. Development of the Colle-Salvetti correlation-energy formula into a functional of the electron density. *Phys Rev B Condens Matter*. 1988;37(2):785–789.
3. Guillemin JC, Wlodarczak G, Lopez JC. Millimeter-wave spectrum of cyclopropanone,  $\text{C}_3\text{H}_2\text{O}$ . *Journal of Molecular Spectroscopy*. 1990;140:190–192.
4. Sharma MK, Sharma M, Chandra S. Suggestion for search of cyclopropanone ( $c\text{-C}_3\text{H}_2\text{O}$ ) in a cosmic object. *Molecular Astrophysics*. 2017;6:1–8.
5. Frisch MJ. *Gaussian Inc., Wallingford CT*. 2010.
6. Sharma MK, Mushrif PG, Sharma M, Chandra S. Suggestion for the search for  $\text{TiH}_2$  molecule in an interstellar molecular cloud. *Astronomische Nachrichten*. 2017;338(6):715–719.
7. Pan J, Albert S, Sastry KVLN, De Lucia FC. The Millimeter- and Submillimeter-Wave Spectrum of Ethylene Oxide ( $c\text{-C}_2\text{H}_4\text{O}$ ). *The Astrophysical Journal*. 1998;499:517–519.
8. Sharma MK, Sharma M, Chandra S. Suggestion for search of ethylene oxide ( $c\text{-C}_2\text{H}_4\text{O}$ ) in a cosmic object. *Astrophysics and Space Science*. 2018;363:94–101.
9. Sharma MK, Sharma M, Chandra S. Strengths of rotational lines from  $\text{H}_2\text{CC}$  molecule: Addressing tentative detection. *Molecular Astrophysics*. 2018;12:20–24.
10. Bailleux S, Bogey M, Demuyneck C. Millimeter-wave rotational spectrum of  $\text{H}_2\text{SiO}$ . *The Journal of Chemical Physics*. 1994;101:2279–2733.
11. Sharma MK, Sharma M, Chandra S. Suggestion for search of silanone ( $\text{H}_2\text{SiO}$ ) in interstellar medium. *Advances in Space Research*. 2018.
12. Winnenwiser M, Winnenwiser BP, Stein M, et al. Rotational Spectra of  $\text{cis-HCOOH}$ ,  $\text{trans-HCOOH}$ , and  $\text{trans-H}^{13}\text{COOH}$ . *Journal of Molecular Spectroscopy*. 2002;216(2):259–265.
13. Sharma MK. Transfer of radiation in the formic acid - a precursor for amino acids. *Advances in Space Research*. 2017.
14. McCarthy MC, Joshua H Baraban, Joshua H Baraban, et al. Discovery of a Missing Link: Detection and Structure of the Elusive Disilicon Carbide Cluster. *Journal of Physical Chemistry Letters*. 2015;6(11):2107–2111.
15. Sharma MK. *The Disilicon carbide ( $\text{Si}_2\text{C}$ ) molecule in interstellar medium. Molecular Astrophysics*. 2017.
16. Sharma MK, Sharma M, Chandra S. Some observations about MOLPRO. *Physics & Astronomy International Journal*. 2018;2(4):286–288.
17. Werner HJ, Knowles PJ. MOLPRO. Abilities program (2008, 2015). 2018. 1–737.
18. Hutson JM, Green S. *MOLSCAT Computer CODE, Version 14 (1994) Distributed by Collaborative Computational Project No. 6*. The Engineering and Physical Sciences Research Council (UK). 1994.
19. Sharma MK, Sharma M, Chandra S. Some Observations about the MOLSCAT. *Communications in Theoretical Physics*. 2015;64:731–734.